

1 **MANTLE HELIUM IN SOUTHERN QUEBEC GROUNDWATER: A POSSIBLE FOSSIL**
2 **RECORD OF THE NEW ENGLAND HOTSPOT**

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16 England hotspot.

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20 **Abstract**

21 The Montereian Hills are an alignment of magmatic intrusions of Cretaceous age located in the
22 St. Lawrence Lowlands, Quebec, Canada. Their origin is controversial and numerous studies have
23 failed to decipher between a hotspot trail or sub-continental magmatism related to the opening of
24 the North Atlantic Ocean. Here, we show that 17.7 ± 9.6 % of the helium of the modern to Holocene-
25 aged groundwater from the regional aquifer is of mantle origin, with a $^3\text{He}/^4\text{He}$ (R) of up to 1.42
26 times the atmospheric ratio (R_a). It suggests that a fossil Montereian Hills magmatic signal,
27 diluted by local radiogenic helium and preserved in the Montereian Hills intrusions, is leached
28 locally by flowing modern or sub-modern groundwater. Helium isotopic measurements by
29 pyrolysis in Montereian Hills bulk rocks and clinopyroxene separates show R/ R_a values of up to
30 4.96, suggesting that fossil mantle helium has been partially preserved in these rocks and their
31 mineral phases. Monte Carlo simulations of a magma aging model shows that the initial $^3\text{He}/^4\text{He}$
32 ratio in these Cretaceous intrusions could have been between $21 \pm 10R_a$ and $33 \pm 28R_a$ (2σ), favoring
33 the hypothesis that the Montereian Hills are the product of the passage of the North American
34 plate over the New England hotspot. This study raises the prospect of using modern groundwater
35 as an archive of mantle He over a hundreds of millions of years timescale.

36

37 **1. Introduction**

38 The Cretaceous Monteregian Hills Igneous Province is a WNW-ESE alignment of nine
39 alkaline intrusions outcropping across the St. Lawrence Lowlands, Quebec, Eastern Canada (Fig.
40 1 and 2b). The magmatic source of these intrusions has been the subject of debate for nearly 40
41 years. Crough (1981) and later Sleep (1990) have suggested that the Monteregian Hills are the
42 result of the passage of the North American plate over the New England hotspot (also called the
43 Great Meteor hotspot), a long-lived hotspot presently overridden by the Mid-Atlantic Ridge, south
44 of the Azores. However, numerous geochemical and isotopic investigations of Monteregian Hills
45 rocks have failed to unequivocally indicate the presence of a plume source (e.g., McHone, 1996;
46 Roulleau et al., 2012; Roulleau and Stevenson, 2013; Chen and Simonetti, 2015).

47 The helium isotope ratio, $^3\text{He}/^4\text{He}$ (R), can be used to trace the sources of continental
48 magmatism. Indeed, the mantle is enriched in primordial isotope ^3He over radiogenic isotope ^4He
49 when compared to the atmosphere ($R_a = 1.382 \times 10^{-6}$; Sano et al., 2013). In continental settings,
50 groundwater may contain helium derived from the mantle in regions of active transtensional
51 deformation, even in regions where there is no substantial magmatism (e.g., Kennedy and van
52 Soest, 1997). The processes that control the release of mantle-derived volatiles to the subsurface
53 are likely sub-continental magma intrusions and underplating (Torgersen, 1993), with episodic
54 advective transport in the crust (Castro, 2004). The occurrence of mantle helium should be
55 identifiable in crustal environments, which are generally dominated by radiogenic ^4He , produced
56 through the radioactive decay of U and Th contained in rocks, over ^3He , which can be produced
57 only in minor amounts by neutron reactions with lithium (Tolstikhin et al., 1996). The resulting
58 typical crustal $^3\text{He}/^4\text{He}$ ratio is 0.01-0.02 R_a (Tolstikhin et al., 1996), distinctly lower than the ratio

59 expected for mantle-related fluids. The helium isotopic ratio of the convective mantle, as sampled
60 in the mid-ocean ridge basalts (MORBs), is $8 \pm 1 R_a$ (e.g., Allègre et al., 1995), while sub-continental
61 lithospheric mantle (SCLM) has a lower ratio, $6.1 \pm 0.9 R_a$ (Gautheron and Moreira, 2002).

62 The non-degassed lower mantle, the expected source of mantle plumes and their related
63 products (Oceanic Island Basalts or OIBs), should have preserved a primordial, solar-type $^3\text{He}/^4\text{He}$
64 ratio, with R/R_a values greater than the MORBs value of 8 ± 1 (e.g., Tieloff et al., 2000). However,
65 measurements from OIBs show highly variable $^3\text{He}/^4\text{He}$ ratios. Some have higher $^3\text{He}/^4\text{He}$ than
66 MORB values, up to $30 R_a$ in Loihi seamount (Hawaii) (Kurz et al., 1982) and greater than $40 R_a$
67 in NW Iceland (e.g., Harðardóttir et al., 2018), but several others have lower values than MORBs,
68 such as Tristan de Cunha and Gough Islands ($5-7 R_a$; Kurz et al., 1982). The highest R/R_a value
69 reported so far is 49.5, measured in picrites from Baffin Island (Stuart et al., 2003). The reason for
70 such high variability is still a matter of debate, and several hypotheses have been proposed,
71 including: 1) ubiquitous air contamination during basalt eruption; 2) the addition of radiogenic
72 helium from subducted plate components; 3) mixing between plume-derived material and material
73 derived from the upper mantle; and 4) isotopic heterogeneity within the plume itself (e.g., Kurz et
74 al., 1982; Graham, 2002).

75 Over the past 25 years, several studies have identified fossil mantle helium in groundwater
76 from the eastern North American craton. In New Hampshire, Torgersen et al. (1995) identified the
77 presence of helium with a substantial mantle fraction ($\approx 12\%$ of the total; $^3\text{He}/^4\text{He} = 1.2 R_a$) in
78 modern groundwater from the Mirror Lake Basin aquifer, where there has been no volcanism since
79 the Jurassic to Early Cretaceous (190-95 Ma ago). Torgersen et al. (1995) attributed this mantle
80 helium to either the passage of the New England hotspot, which created the nearby White

81 Mountains magmatic province (Eby, 1985) or to magmatic episodes related to the closure of the
82 Iapetus ocean (before 370 Ma). Castro et al. (2009) measured He and Ne in deep groundwater of
83 the Michigan Basin revealing a primordial solar-like contribution. However, the absence of
84 magmatism and of low-velocity seismic anomalies that would have indicated a fossil hotspot track
85 in the region suggested that this primordial signature could be accounted for by a shallow refractory
86 mantle reservoir underneath the Michigan Basin (Castro et al., 2009). Pinti et al. (2011) identified
87 small amounts of mantle helium (2% of the total) in hundreds of millions of years-old brines located
88 in the deeper sedimentary aquifers of the St. Lawrence Lowlands (BEC location in Fig. 1). Saby et
89 al. (2016) also found mantle helium in the Ordovician-age regional aquifer of the Nicolet St-
90 François watershed, St. Lawrence Lowlands (8% of the total; NSF location in Fig. 1).

91 Here, three new sets of helium isotopic data from groundwater and rocks collected from
92 several locations in the St. Lawrence Lowlands are reported (Figs. 2a, 2b). The first set was
93 measured in groundwater collected from sedimentary aquifers of the Vaudreuil-Soulanges
94 watershed (Table 1), located west of Montreal, close to the Oka carbonatite complex (Fig. 2a). The
95 second set of helium data was measured in groundwater circulating in the magmatic intrusions of
96 the Monteregian Hills (Table 1 and Fig. 2b). The third was measured in Monteregian Hills bulk
97 rocks and clinopyroxene separates, collected as part of a previous study (Rouilleau and Stevenson,
98 2013; Fig. 2b). Measured $^3\text{He}/^4\text{He}$ ratios in both groundwater and rocks suggest the presence of a
99 fossil helium magmatic signal diluted by local radiogenic helium, preserved in the Monteregian
100 intrusions and leached locally by flowing modern to Quaternary-age groundwater. The initial
101 helium signature of the magmatic source, before radiogenic dilution, is estimated here using Monte
102 Carlo simulations of a magma aging model (Torgersen and Jenkins, 1982). The results bring new

103 and independent evidence of the past presence of a primitive plume beneath the region, which
104 generated the Montereian intrusions.

105

106 ***2. Geological background***

107 *2.1 Hydrology of sampled waters*

108 The first set of groundwater samples was collected 10 km south of the Oka complex, in the
109 Vaudreuil-Soulanges watershed (Table 1 and Fig. 2a). With a surface area of 814 km², Vaudreuil-
110 Soulanges is the westernmost watershed of the St. Lawrence Lowlands basin (Fig. 2a). It consists
111 of Cambrian fluvio-marine quartzitic sandstones of the Potsdam Group and Early Ordovician shelf-
112 carbonates and dolostones of the Beekmantown Group unconformably overlying the granite-
113 gneiss-anorthosite basement of the Mesoproterozoic (1250-980 Ma) Grenville Orogeny (Larocque
114 et al., 2015). This sedimentary sequence constitutes the bedrock aquifer sampled in this study, and
115 is covered by a 120 m-thick sequence of Quaternary glacial till and clays from the marine
116 transgression episode of the Champlain Sea, which occurred between 12 and 9 ka ago.

117 The sedimentary sequence of the Vaudreuil-Soulanges watershed is disturbed by two main
118 magmatic episodes: (1) the intrusion of the syeno-granitic body of Mont Rigaud, with a U-Pb age
119 of 564_{-8}^{+10} Ma (Malka et al., 2000); and 2) K-rich alnoïtes intruding the sedimentary cover in the
120 central-eastern part of the watershed at the location of Île Cadieux, with ages ranging 108 to 113
121 Ma (Fig. 2a; Eby, 1987; Chen and Simonetti, 2013). Although only the latter intrusion outcrops in
122 the watershed, a regional low-resolution aeromagnetic survey of the St. Lawrence Lowlands has
123 shown that hundreds of smaller intrusions lie under the sedimentary cover, with a high magnetic

124 alignment corresponding to the northern region of the Vaudreuil-Soulanges watershed (see maps
125 at <http://www.mrnfp.gouv.qc.ca/produits-services/mines.jsp>). Several of these intrusions outcrop
126 in numerous quarries, or reach the surface, by crosscutting the Ordovician sedimentary sequences
127 of the St. Lawrence Lowlands (e.g., Harnois et al., 1990; Harnois and Mineau, 1991).

128 Recharge of the Cambrian-Ordovician bedrock aquifer is mainly from high-permeability
129 Quaternary eskers and sand hills, which crosscut the low-permeability glacial till and the
130 Champlain Sea clays. Less recharge is from the higher topography zone, corresponding to the Mont
131 Rigaud syenitic intrusion. Regional groundwater flows from these topographic highs to the Ottawa
132 River in the northern part of the area, and S-SE, in the direction of St. Lawrence River and Lake
133 Saint-Francois in the southern part of the basin (Fig. 2a). Méjean (2016) reported modelled ^3H - ^3He
134 groundwater ages to be from 8.5 ± 0.6 a to 57 ± 4 a, with several samples showing ^3H of 0.8 Tritium
135 Units (TU; 1 TU is defined as the ratio of 1 tritium atom to 10^{18} hydrogen atoms), which
136 corresponds to the environmental background, suggesting the presence of older water masses. This
137 fossil water component was dated using the U-Th/ ^4He method to be older than 45 ka (Méjean,
138 2016).

139 The second set of groundwater samples was collected from wells tapping directly into the
140 fractured intrusive rocks of the Monteregian Hills at St. Bruno, St. Hilaire, Rougemont, Yamaska,
141 and Brome (Fig. 2b and Table 1). Little is known about groundwater circulation in the Monteregian
142 Hills. Geophysics have revealed the presence of several buried dykes around these hills, but the
143 consequences for groundwater circulation are still unknown. A recent study by Beaudry et al.
144 (2018) revealed that the Monteregian Hills contain confined to semi-confined aquifers, with water
145 ages between modern and 3-12 ka. Groundwater from the Monteregian Hills belongs to two

146 chemical groups (Beaudry et al., 2018): 1) Group “A1”, which consists of modern recharge water
147 associated with the presence of tritium, low salinity (Total Dissolved Salinity (TDS) of 168 mg/L),
148 and a Ca-HCO₃ water type with relatively low pH (6.5) and a composition controlled by carbonate
149 dissolution; and 2) group “A2”, which consists of evolved Na-HCO₃ composition water with high
150 pH (7.8) and slightly saline (TDS of 368 mg/L) due to mixing with brackish water of the Champlain
151 Sea marine invasion (Beaudry et al., 2018).

152

153 *2.2 The debated origin of the Montereian Hills*

154 The Montereian Hills alkaline province consists of plutons, dikes, and sills that were
155 emplaced along faults associated with the Ottawa-Bonnechere graben (Fig. 2). The intrusions were
156 emplaced into the Cambro-Ordovician sediments of the St. Lawrence Lowlands and the folded and
157 thrust Lower Paleozoic sequence of the Appalachian orogen (Figs. 1 and 2b). Carbonatite and
158 associated alkaline rocks occur at the Oka complex, just north of the Vaudreuil-Soulanges
159 watershed (Fig. 2a). Further east, the plutons are dominated by ultramafic and mafic rocks
160 (gabbros) with fewer syenites. Toward the eastern end of the province, silica-saturated mafic and
161 felsic rocks become dominant components of the plutons. Geochronological data from the Oka
162 carbonatite complex by Chen and Simonetti (2013) show bimodal U-Pb ages centered on 113 Ma
163 for alnöitic intrusions and 135 Ma for ijolite rocks. Measured ⁴⁰Ar/³⁹Ar ages of the other
164 Montereian complexes fall within the restricted range of 124±1 Ma (Foland et al., 1986).

165 The alkaline character of the Montereian Hills and the Sr (0.7032–0.7040) and Nd
166 (0.5125–0.5127) isotopic composition similar to those found in some OIBs supported a plume-
167 related mantle source (Foland et al., 1986). McHone (1996) showed that there is no evidence of

168 age progression within and among the different magmatic provinces supposedly created by the
169 New England hotspot (Monteregian Hills in Quebec, White Mountains in New England, and the
170 New England seamounts off the east coast of the US), as would be expected of a hotspot track,
171 arguing for the involvement of heterogeneous SCLM melts rather than plume material. A decrease
172 in the buoyancy flux of the New England hotspot from the New England seamounts to the
173 Monteregian Hills was estimated by Sleep (1990), who explained that the lack of an obvious track
174 may be due to either the failure of the plume to penetrate the Canadian Shield. Roulleau et al.
175 (2012) observed an inverse relationship between $^{206}\text{Pb}/^{204}\text{Pb}$ and $\text{N}_2/^{36}\text{Ar}$ ratios, interpreted as the
176 mixing between a plume source with a recycled component (HIMU, High $\mu = ^{238}\text{U}/^{204}\text{Pb}$) and an
177 ambiguous mantle source that could be either a non-HIMU-type plume source (such as that feeding
178 Loihi; Dixon and Clague, 2001) or the depleted mantle. Roulleau and Stevenson (2013) provided
179 new isotopic data (Nd-Sr-Hf-Pb) suggesting a dominant lithospheric mantle source for the
180 Monteregian Hills, but contaminated by rising asthenospheric mantle that progressively
181 thermomechanically eroded the lithosphere, hence, producing a metasomatized lithospheric mantle
182 with HIMU and EMI reservoirs. Finally, some other authors have ascribed the Monteregian
183 magmatism to the melting of the lithospheric mantle through NW-SE Iapetus-related faults
184 reactivated at the time of the North Atlantic Ocean opening (e.g., Faure et al., 1996).

185

186 **3. Materials and analytical methods**

187 *3.1 Groundwater sampling*

188 A total of 24 groundwater samples were collected from open bedrock municipal wells,
189 domestic wells, and piezometers. Sixteen were sampled in the Cambrian-Ordovician fractured

190 bedrock aquifer of the Vaudreuil-Soulanges watershed in the summer of 2014, at depths ranging
191 between 20 m and 96 m. A second set of eight groundwater samples was collected in the summer
192 of 2016, from fractured aquifers of the Monteregian Hills: Mount St. Bruno, Mount St. Hilaire,
193 Mount Rougemont, Mount Yamaska and Mount Brome (Fig. 2b and Table 1). After purging the
194 well, water was collected in 3/8-inch diameter refrigeration-type copper tubes, which were then
195 cold-sealed following the method described in Vautour et al. (2015). Helium isotopic ratios of
196 groundwater samples collected in 2014 were measured at the University of Tokyo, using a VG
197 Helix SFT and compared to the Helium Standard of Japan (HESJ; Matsuda et al., 2002), with a 2σ
198 precision of $\pm 1\%$. The ^4He and ^{20}Ne concentrations were measured on a Pfeiffer Prisma™ 80
199 quadrupole mass spectrometer connected to the purification line (see Vautour et al., 2015 for
200 details).

201 The eight samples collected in 2016 were analyzed at the GEOTOP Noble Gas Laboratory.
202 After extraction and purification on one Ti-sponge and two ST-707 getters, helium was
203 cryogenically separated and released at 40K into a HELIX-MC (Thermo®) mass spectrometer, and
204 $^3\text{He}/^4\text{He}$ ratios were measured by peak jumping on the axial CDD (^3He) and Faraday (^4He)
205 detectors, with a precision of ca. 3% on the ^3He and less than 0.01% on ^4He , calibrated against an
206 air pipette. Helium and neon elemental composition was measured in a Pfeiffer Prisma™ C-200
207 quadrupole mass spectrometer connected to the purification line, following the procedures
208 described in Roulleau et al. (2012).

209 *3.2 Bulk rocks and mineral separates*

210 A total of nine rock samples (Figs. 2a, b) were collected from the Oka complex (OC; n=1),
211 Île Cadieux (Aln, n=1), Mount Royal (MR; n=1), Mount St. Bruno (MSB, n=1), Mount Rougemont
212 (MR, n=2), Mount Brome (MB, n=1), Mount St. Hilaire (MSH, n=1), and Mount Yamaska (MY,
213 n=1). Pyroxene is among the first minerals to crystallize in the magma chamber, together with
214 olivine, and thus documents the least evolved melt available. Montereian Hills gabbros contain
215 very small and few olivine crystals, so clinopyroxene remains the main phase to analyze (Sasada
216 et al., 1997; Roulleau et al., 2012). Clinopyroxene grains were thus hand-picked from the Mount
217 St. Bruno sample (SMB3-R; Table 2). Helium was extracted and analyzed at the University of
218 Tokyo using a two-stage analytical process. Whole-rock samples and mineral separates (~0.5 g)
219 were wrapped in 10 µm-thick aluminum foil and stored inside the glass part of an induction furnace
220 under vacuum to degas weakly-bound atmospheric gases. Individual samples were then moved into
221 a Mo-Ta furnace and pyrolyzed at 1800°C for 30 mn (Kagoshima et al. 2012). The gas was purified,
222 cryogenically separated, and He contents were determined on a Pfeiffer QMS Prisma™ quadrupole
223 mass spectrometer connected to the vacuum purification line. The ³He/⁴He ratios were measured
224 using the HELIX-SFT mass spectrometer (R-OC and R-Aln) and the VG5400 mass spectrometer
225 (R-MR15, R-MSB3, R-MB15, R-MRg1, R-MRg9, R-MSH1, R-MY3) following the procedures
226 described above.

227
228 The rock samples collected in the Vaudreuil-Soulanges and the Montereian Hills areas
229 were analyzed by Bureau Veritas® laboratories in Vancouver for trace elements, including U and
230 Th, reported in Table 2. Samples were finely crushed and mixed with LiBO₂/Li₂B₄O₇ flux.
231 Crucibles were fused in a furnace. The cooled bead was dissolved in ACS grade nitric acid and

232 trace elements were analyzed using two different instruments: Elan 9000 ICP-MS and Spectro
233 Ciros Vision ICP-ES. Loss on ignition (LOI) was determined by igniting a sample split then
234 measuring the weight loss. The detection limits for U and Th by ICP-MS are 0.1 and 0.2 ppm
235 respectively.

236

237 **4. Results**

238 The atmosphere-normalized $^3\text{He}/^4\text{He}$ ratios (R/R_a) of the groundwater samples range from
239 0.15 ± 0.01 to 3.32 ± 0.08 , pointing to the presence of both ^3He -enriched and ^4He -enriched sources
240 (Table 1). Dissolved ^4He concentrations vary over three orders of magnitude. The lowest helium
241 concentration is $6.64 \times 10^{-8} \text{ cm}^3\text{STP/g}_{\text{H}_2\text{O}}$ (VS104), which is slightly higher than the value expected
242 from solubility equilibrium with the atmosphere (Air Saturated Water or ASW) at 9.8°C , of $4.7 \times$
243 $10^{-8} \text{ cm}^3\text{STP/g}_{\text{H}_2\text{O}}$, which represents helium contained in freshwater at recharge. Two samples show
244 very high ^4He concentrations: Mount St. Hilaire (WMSH3), with ^4He of $0.0786 \text{ cm}^3\text{STP/g}_{\text{H}_2\text{O}}$, and
245 Mount Brome (WMB8), with ^4He of $1.27 \times 10^{-4} \text{ cm}^3\text{STP/g}_{\text{H}_2\text{O}}$. These high ^4He concentrations are
246 also accompanied by high ^{20}Ne concentrations. In the case of WMSH3, the high $^4\text{He}/^{20}\text{Ne}$ of 49.80
247 excludes the presence of excess air and the contemporary He-Ne enrichment could be related to
248 bubbling and a bi-phase (gas-water) mixture not sampled at the solubility equilibrium, as observed
249 in a few wells in the St. Lawrence Lowlands (Vautour et al., 2015). The ^4He concentrations have
250 been corrected assuming a ^{20}Ne content at solubility equilibrium of 9.3×10^{-7} and 1.65×10^{-8}
251 $\text{cm}^3\text{STP/g}_{\text{H}_2\text{O}}$ for WSMH3 and WMB8, respectively. Sample WMSH3 also shows the lowest and
252 most radiogenic $^3\text{He}/^4\text{He}$ ratio, of 0.15 ± 0.01 , which is caused by the high U and Th contents in this

253 Monteregian intrusion. Mount St. Hilaire is indeed known for its high natural radioactivity, which
254 produces high ^{222}Rn concentrations in groundwater (234 Bq/L; Pinti et al., 2014).

255 The $^3\text{He}/^4\text{He}$ ratios measured in the bulk rocks range from 0.06 ± 0.02 (R-Aln) to
256 $1.36\pm 0.11\text{Ra}$ (R-Mrg9). Higher values were obtained in the clinopyroxene separated from sample
257 R-MSB3, with $\text{R/Ra} = 1.23\pm 0.10$ and 4.96 ± 0.01 . The total amount of ^4He measured ranges from
258 $7.90 \times 10^{-7} \text{ cm}^3\text{STP/g}_{\text{rock}}$ to $2.28 \times 10^{-5} \text{ cm}^3\text{STP/g}_{\text{rock}}$ (Table 2). The measured $^3\text{He}/^4\text{He}$ ratios in
259 mineral separates is similar to, or higher than, those measured by Sasada et al. (1997) in Oka
260 intrusion diopside, of 0.25 ± 0.07 and $2.55\pm 0.25\text{Ra}$ (Table 2).

261

262 **5. Discussion**

263 *5.1 Helium sources in Vaudreuil-Soulanges and Monteregian Hills groundwater*

264 Helium dissolved in groundwater can have multiple sources. Each helium component must
265 be correctly identified in order to determine the mantle-derived one, if any. In groundwater, the
266 helium components are:

$$267 \quad {}^3\text{He}_{\text{tot}} = {}^3\text{He}_{\text{eq}} + {}^3\text{He}_{\text{tri}} + {}^3\text{He}_{\text{ea}} + {}^3\text{He}_{\text{rad}} + {}^3\text{He}_{\text{mtl}} \quad (1) \text{ and}$$

$$268 \quad {}^4\text{He}_{\text{tot}} = {}^4\text{He}_{\text{eq}} + {}^4\text{He}_{\text{ea}} + {}^4\text{He}_{\text{rad}} + {}^4\text{He}_{\text{mtl}} \quad (2),$$

269 where the suffix “eq” indicates atmospheric helium dissolved in groundwater during recharge at
270 solubility equilibrium (ASW); “ea” is the excess-air helium resulting from air bubbles entering the
271 water table at recharge and dissolving into groundwater (Heaton and Vogel, 1981); “tri” is the ^3He

272 produced by the decay of atmospheric tritium dissolved in water, commonly referred to as
273 tritiogenic helium; “mtl” is the mantle helium; and “rad” is the radiogenic helium produced in the
274 aquifer. The latter can produce ^3He from neutron reactions on lithium ($^6\text{Li}(n,\alpha)^3\text{H}(\beta)^3\text{He}$) and ^4He
275 are the α -particles supplied through the decay of $^{235,238}\text{U}$ and ^{232}Th and their decay products
276 (Andrews and Kay, 1982). Radiogenic and mantle helium are commonly grouped using the term
277 “terrigenic”.

278 There is another potential source of ^3He , which is often ignored: cosmogenic ^3He produced
279 by spallation reactions between cosmic ray neutrons and the major elements of the rock.
280 Cosmogenic ^3He is produced in rocks in the uppermost 1 m of Earth’s surface and could potentially
281 be transferred to groundwater. However, calculations from Marty et al. (1993) of potential
282 cosmogenic ^3He in groundwater of the Paris Basin have shown that cosmogenic production in rocks
283 is very limited as a potential source, and cannot account for significant ^3He anomalies in
284 groundwater. Sediments from Vaudreuil-Soulanges watershed outcropped in Ordovician time (430
285 Ma ago), but were then rapidly buried for several hundred million years, so that any cosmogenic
286 ^3He produced is expected to have been completely diluted by continuous radiogenic ^4He
287 production, and thus not detectable in our samples.

288 A convenient graphical approach to separate the different helium components is to plot the
289 $^3\text{He}/^4\text{He}$ ratio corrected for the excess-air component ($[(^3\text{He}_{\text{tot}} - ^3\text{He}_{\text{ea}})]/[^4\text{He}_{\text{tot}} - ^4\text{He}_{\text{ea}}]$) against the ratio
290 of ASW helium (He_{eq}) to total helium, corrected for the excess-air component ($^4\text{He}_{\text{eq}}/[^4\text{He}_{\text{tot}} - ^4\text{He}_{\text{ea}}]$)
291 (Weise and Moser, 1987). In our samples, the excess air component is calculated as the difference
292 between the expected ASW value of ^{20}Ne at the well temperature and the measured one.

293 In this diagram, which is commonly referred to as a “Weise plot” (Fig. 4a), helium of
 294 purely atmospheric origin (ASW composition) plots on the coordinates ($[^3\text{He}_{\text{tot}}-^3\text{He}_{\text{ea}}]/[^4\text{He}_{\text{tot}}-^4\text{He}_{\text{ea}}]$)
 295 or $R_{\text{eq}} = 1.36 \times 10^{-6}$ (0.983Ra; Benson and Krause, 1980) and $(^4\text{He}_{\text{eq}}/^4\text{He}_{\text{tot}}-^4\text{He}_{\text{ea}}) = 1$ respectively
 296 (Fig. 4a). If only tritiogenic $^3\text{He}_{\text{tri}}$ is added to an initial atmospheric composition, then the sample
 297 composition will plot parallel to the right Y-axis, starting from the initial ASW composition (Fig.
 298 4a). If radiogenic ^3He , produced through neutron reactions with ^6Li , and ^4He , produced through the
 299 α -decay of $^{235,238}\text{U}$ and ^{232}Th , are added, the initial ASW composition will shift closer to the bottom-
 300 left corner of the plot (R_{rad} ; Fig. 4a). In the St. Lawrence Lowlands fractured bedrock, the crustal
 301 R_{rad} is calculated to be 1.66×10^{-6} (0.012Ra), based on Li, U, and Th contents measured from 398
 302 rocks of the region (Pinti et al., 2011). Finally, if mantle-derived helium is present, sample
 303 composition will plot parallel to the Y-axis, starting from a radiogenic crustal composition (R_{terr} ;
 304 Fig. 4a).

305 On the Weise plot, mixing between helium components appears as straight lines governed
 306 by the equation (Weise and Moser, 1987; Fig. 4a):

$$307 \quad \underbrace{\left(\frac{^3\text{He}_{\text{tot}}-^3\text{He}_{\text{ea}}}{^4\text{He}_{\text{tot}}-^4\text{He}_{\text{ea}}}\right)}_Y = \underbrace{\left(R_{\text{eq}} - R_{\text{terr}} + \frac{^3\text{He}_{\text{tri}}}{^4\text{He}_{\text{eq}}}\right)}_m \cdot \underbrace{\left(\frac{^4\text{He}_{\text{eq}}}{^4\text{He}_{\text{tot}}-^4\text{He}_{\text{ea}}}\right)}_X + \underbrace{R_{\text{terr}}}_b \quad (3),$$

308 where R_{terr} is the $^3\text{He}/^4\text{He}$ of the terrigenous source, which corresponds to either crustal helium,
 309 mantle-derived helium, or both.

310 Groundwater samples from Vaudreuil-Soulanges and the Monteregian Hills align along
 311 mixing lines representing the complete decay of 6.7TU, 14.5TU, 18.5TU, 40TU, 61TU, and
 312 285TU, and a terrigenous end-member, R_{terr} , with values between 0.16Ra and 1.48Ra ($^3\text{He}/^4\text{He}$ values

313 between 2.2×10^{-7} and 2.05×10^{-6} ; Fig. 4b). These ratios are higher than those expected from the
314 local production of ^3He from ^6Li and of ^4He from $^{238,235}\text{U}$ and ^{232}Th in St. Lawrence Lowlands
315 sedimentary rocks or in rocks from the Grenville basement, of 0.45 to 1.66×10^{-8} (Pinti et al.,
316 2011). The difference between the expected crustal R_{rad} and the maximum calculated R_{terr} from the
317 Weise plot (i.e., the intercept of the mixing line between samples WSB1, WSB2, and WY6; Fig.
318 4b) suggests the addition of $18.4 \pm 6.7\%$ of mantle-derived helium (labelled “ f_{mtl} ” in Fig. 4b). The
319 total uncertainty on the value comes mainly from the error on the intercept of the regression.
320 Uncertainty derived from the variability of the mantle end-member helium isotopic ratio (8 ± 1) and
321 the possible variability on the crustal ratio (0.01 to $0.1R_a$; Tolstikhin et al., 1996) produces less
322 than 2% error. It is worth noting that here a $^3\text{He}/^4\text{He}$ mantle end-member ratio of 8 ± 1 is assumed
323 for sake of consistency with previous estimates from Pinti et al. (2011) and Saby et al. (2016). If
324 the closer mixing line between samples VS-105-2, VS-108-2, and VS-108-2 and the sample from
325 Mount Rougemont (WR-5) is taken, the intercept will give a R_{terr} of 1.86×10^{-6} (or $1.35R_a$; Fig.
326 4b). This is equivalent to a mantle fraction of $16.7 \pm 6.9\%$. This value is, within uncertainty, similar
327 to the previous estimate.

328 These two f_{mtl} values represent the highest mantle helium fractions ever reported for St.
329 Lawrence Lowlands groundwater, higher than the 2% and 8% calculated by Pinti et al. (2011) and
330 Saby et al. (2016), respectively. However, the calculated mantle He fractions could be conservative
331 minimal estimates. The complete decay of 1TU in a gram of water corresponds to 2.49×10^{-15}
332 $\text{cm}^3\text{STP}/\text{g}_{\text{water}}$ of ^3He . Assuming an initial $^4\text{He}_{\text{ASW}}$ of $4.72 \times 10^{-8} \text{cm}^3\text{STP}/\text{g}_{\text{water}}$, corresponding to the
333 dissolution of atmospheric helium at 9.8°C , which is the average temperature of the groundwater
334 in the region (Vautour et al., 2015), the corresponding $^3\text{He}_{\text{ASW}}$ would be $6.42 \times 10^{-14} \text{cm}^3\text{STP}/\text{g}_{\text{water}}$.

335 Adding 285TU of tritium would produce a tritiogenic $^3\text{He}/^4\text{He}$ ratio of 11.86Ra. This value is not
336 unreasonable for aquifers that contain a mixture of modern and older water, but is actually more
337 than double the highest tritiogenic $^3\text{He}/^4\text{He}$ ratio measured among the 125 groundwater samples
338 from Bécancour and Nicolet-St. François watersheds (internal compilation of unpublished data),
339 and equal to $4.59 \pm 0.03\text{Ra}$ (equivalent to the decay of 94.5TU). If the maximum amount of tritium
340 preserved in groundwater of the St. Lawrence Lowlands aquifers is 94.5TU, then the mixing line
341 passing through, e.g., sample WSB1 (mixing line reported as an example in Fig. 4a), will have an
342 intercept (R_{terr}) at $^3\text{He}/^4\text{He} = 4.28 \times 10^{-6}$, which is equivalent to the addition of 38.6% of mantle
343 helium. For the purpose of this study and the following modeling, the average value of the two f_{mtl}
344 estimates above ($18.4 \pm 6.7\%$ and $16.7 \pm 6.9\%$), i.e. $17.7 \pm 9.6\%$ (equivalent to a $^3\text{He}/^4\text{He}$ ratio of
345 $1.42 \pm 0.14\text{Ra}$), is assumed here (orange star in Fig. 4b).

346

347 *5.2 Origin of terrigenic helium in Vaudreuil-Soulanges and Monteregian Hills groundwater*

348 The ^3He in excess of crustal production (Fig. 4b) can be explained by the presence of a
349 mantle helium component in groundwater of the Vaudreuil-Soulanges watershed and in the
350 Monteregian Hills local aquifers. The preservation of this mantle component in groundwater could
351 be explained by 1) mantle He concentrated at the base of the crust and migrating through crustal-
352 rooted faults; 2) a fossil mantle source preserved in very old brines having ages contemporary with
353 the emplacement of the Monteregian Hills intrusions; or 3) a fossil mantle signal preserved in
354 Monteregian Hills magmatic intrusions, but released in recent groundwater by water-rock
355 processes, such as mineral dissolution or diffusive exchange.

356 The first hypothesis can be discarded. For example, Lee et al. (2019) have shown that
357 mantle He can reach shallow groundwater through active strike-slip faults. Fluids are thought to
358 enhance pore fluid pressures in association with fault weakening caused by seismicity. This will
359 facilitate the upward migration of deep-seated helium. Although Quebec is the second more seismic
360 area in Canada, seismicity seems to be caused by residual friction provoked along fault planes by
361 the crustal isostatic rebound following the last deglaciation (Lamontagne, 2002). Fission-track and
362 U-Th/He thermochronology of the Cambrian St-Lawrence Lowlands Rift Zone showed that fault
363 activity in the region continued until the Mesozoic and terminated in the Cretaceous (Tremblay et
364 al., 2013). Consequently, it is difficult to assume that present-day micro-seismicity has weakened
365 these old tectonic structures to facilitate the upward migration of mantle helium in the region.

366 If the second hypothesis is valid, then old groundwater of at least the age of the Montereian
367 Hills intrusions (124 ± 1 Ma) should have been preserved in the study area. Pinti et al. (2011) found
368 brines with possible Mesozoic ages in the Cambro-Ordovician aquifers of the Potsdam Group of
369 the St. Lawrence Lowlands supergroup in the Bécancour area (Fig. 1). However, these brines have
370 little mantle-derived ^3He (2% of the total), are highly saline Ca-Na-Cl type (up to 315,000 mg/L)
371 waters, and are very radiogenic ($^{87}\text{Sr}/^{86}\text{Sr}$ of 0.71214), characteristic of Precambrian basement
372 brines. Groundwater from Vaudreuil-Soulanges and from Montereian Hills aquifers are of Ca-
373 HCO_3 or Na- HCO_3 types, with very low salinity (168-710 mg/L), mainly derived from exchange
374 with glacio-marine porewater from the Champlain Sea clays (Larocque et al., 2015; Beaudry et al.,
375 2018). The presence of tritium in several Vaudreuil-Soulanges groundwater samples suggests
376 modern water, possibly mixed with Holocene to Quaternary glacial water containing
377 predominantly radiogenic ^4He , as suggested by Méjean (2016). Similar ages, from modern to 13

378 ka old, have been proposed for Montereian Hills groundwater (Beaudry et al., 2018). Therefore,
379 the occurrence of Mesozoic brines in the Vaudreuil-Soulanges and Montereian Hills aquifers
380 appears to be unlikely.

381 The third alternative and most plausible explanation, which will be explored in detail in the
382 following sections, is the presence of a fossil mantle source, diluted over time through the addition
383 of local radiogenic ^4He (the so-called “magma aging model”; Torgersen and Jenkins, 1982),
384 preserved in magmatic intrusions, and recently released into groundwater through leaching.

385 *5.3 Impact of in situ helium production on the “magma aging” model*

386 The ^3He production in the crust is controlled by the thermal neutron capture of ^6Li in the
387 reaction $^6\text{Li}(n,\alpha)^3\text{H}(\beta^-)^3\text{He}$ (Andrews and Kay, 1982), while ^4He production is regulated by the α -
388 decay of ^{238}U , ^{235}U , and ^{232}Th . In this study, the nucleogenic production rate of ^3He was calculated
389 based on the Li abundances reported by Pinti et al. (2011) and Gold (1966) for the main
390 Montereian Hills lithologies. These vary from 11 ppm for a pyroxenite to 28 ppm for a gabbro
391 (representing the lithologies of samples R-MSB3, R-MB15, R-MRg1, R-MRg9, R-MSH10, and
392 R-MY3), 37 ppm in diorite (sample R-MR-15), and 15 ppm in carbonatite and in alnoïte. U and
393 Th concentrations have been measured in bulk rocks and range from 3.3 to 12.4 ppm and from 9.4
394 to 40.9 ppm respectively (Table 2; data from this study and Roulleau and Stevenson, 2013). Li, U,
395 and Th concentrations were not measurable in mineral separates.

396 The ^3He production rate for Montereian Hills bulk rocks was calculated using the modified
397 equation of Castro (2004):

398 $J_{3\text{He}} = (2.69 \times 10^{-4}[\text{U}] + 6.40 \times 10^{-5}[\text{Th}]) \times [\text{Li}] \times 10^{-23} \times 22414$ (4),

399 where $J_{3\text{He}}$ is the ^3He production rate in $\text{cm}^3\text{STP/g}_{\text{rock}}/\text{yr}$; [U], [Th] and [Li] are the U, Th and Li
400 concentrations in ppm. The ^3He production rate varies from $0.63 \times 10^{-20} \text{cm}^3\text{STP/g}_{\text{rock}}/\text{yr}$ in the Oka
401 carbonatite to $2.89 \times 10^{-20} \text{cm}^3\text{STP/g}_{\text{rock}}/\text{yr}$ in Mont Brome gabbro (Table 2).

402 The ^4He production rate ($J_{4\text{He}}$ in $\text{cm}^3\text{STP/g}_{\text{rock}}/\text{yr}$) is calculated following the equation
403 (Torgersen et al., 1995):

404 $J_{4\text{He}} = (0.2355 \times 10^{-12}) \times [\text{U}] \times \left(1 + 0.123 \left(\frac{[\text{Th}]}{[\text{U}]} - 4\right)\right)$ (5),

405 where [U] and [Th] are the concentrations of U and Th in the rock in ppm (Table 2). The ^4He
406 production rate varies from $0.63 \times 10^{-13} \text{cm}^3\text{STP/g}_{\text{rock}}/\text{yr}$ in the Oka carbonatite and Mount Royal
407 gabbro to $2.07 \times 10^{-12} \text{cm}^3\text{STP/g}_{\text{rock}}/\text{yr}$ in Mount Brome gabbro (Table 2). The resulting $^3\text{He}/^4\text{He}$
408 derived by the *in situ* production of helium isotopes in rocks ranges from 0.005 to 0.0013 (Table
409 2), well below the values measured in fluids, confirming that the groundwater contains a fraction
410 of terrigenic helium derived from a mantle reservoir (Table 1).

411

412 5.4. Monte Carlo simulation of a helium magma aging model

413 To test whether the high terrigenic $^3\text{He}/^4\text{He}$ ratios measured in groundwater (Table 1 and
414 Fig. 4b) could reflect the presence of a fossil mantle source, diluted over time by the addition of *in*
415 *situ*-produced radiogenic ^4He , a “magma aging” model is developed here (Torgersen and Jenkins,
416 1982). This model describes the evolution of an initial magmatic $^3\text{He}/^4\text{He}$ ratio in a magma body
417 since its crystallization, by the addition of radiogenic ^4He produced *in situ* by ^{235}U , ^{238}U , and ^{232}Th
418 α -decay (Torgersen and Jenkins, 1982). Consequently, the U and Th amounts in the magma source

419 determine the rate of change of the helium concentration of the rock, while the initial concentration
 420 of helium in an emplaced magma determines the change in the isotope ratio of helium (R_{mtl}/Ra).
 421 The model does not take the nucleogenic production of ^3He from Li into account, which is
 422 considered here to be negligible.

423 The initial $^3\text{He}/^4\text{He}$ ratio in the magma source (R_{mtl}/Ra) can be calculated following the
 424 equation:

$$425 \quad \left(\frac{R_{\text{mtl}}}{Ra}\right)_{\text{initial}} = \frac{\left[\left(\frac{R_{\text{mtl}}}{Ra}\right)_{\text{final}} \times ({}^4\text{He}_{\text{initial}} + J_{4\text{He}} \times t)\right]}{{}^4\text{He}_{\text{initial}}} \quad (6),$$

426
 427 where $(R_{\text{mtl}}/Ra)_{\text{final}}$ is the $^3\text{He}/^4\text{He}$ ratio measured in modern groundwater and normalized to the air
 428 value; ${}^4\text{He}_{\text{initial}}$ is the amount of ^4He in the magma source at $t=0$; J_{He} is the radiogenic ^4He production
 429 rate as calculated from eqn. (5); and t is the time since the magma emplacement.

430 From the results plotted on the Weise diagram, an average $(R_{\text{mtl}}/Ra)_{\text{final}}$ value of 1.42 ± 0.14
 431 Ra is assumed, equivalent to a mantle fraction f_{mtl} of $17.7 \pm 9.6\%$ (orange star in Fig. 4b). Time, t ,
 432 is assumed to be equal to the average age of alnöite in Vaudreuil-Soulanges ($110.5 \pm 3.5\text{Ma}$), which
 433 ranges from $108 \pm 1\text{Ma}$ and $113 \pm 1\text{Ma}$ (Eby, 1985; Chen and Simonetti, 2013) and of the
 434 Montereian Hills magmatic bodies ($124 \pm 1\text{Ma}$; Foland et al., 1986). The $J_{4\text{He}}$ for the alnöite is 8.41
 435 $\times 10^{-13} \text{ cm}^3\text{STP/g}_{\text{rock}}/\text{yr}$. An uncertainty of 20% is assumed, based on the variability of trace
 436 elements measured in alnöite from Île Cadieux by Harnois and Mineau (1991). The $J_{4\text{He}}$ calculated
 437 for the Montereian Hills is highly variable, depending on the U and Th contents in different
 438 lithologies, from gabbro to syenite (Table 2). Here, an average $J_{4\text{He}}$ of $11.78 \pm 5.01 \times 10^{-13}$
 439 $\text{cm}^3\text{STP/g}_{\text{rock}}/\text{yr}$ is assumed.

440 A critical parameter to determine is the ${}^4\text{He}_{\text{initial}}$. In the magma aging model to explain mantle
441 helium anomalies in groundwater of the Eastern America margin, Torgersen et al. (1995) suggested
442 the value of $6.7 \times 10^{-6} \text{ cm}^3\text{STP/g}_{\text{rock}}$. This value would best represent a gas-rich magma that has
443 retained significant volatiles. The highest ${}^4\text{He}$ concentration in OIBs glasses (database from
444 Graham et al., 2002; n=107) is of $6.24 \times 10^{-6} \text{ cm}^3\text{STP/g}_{\text{rock}}$, very close to the theoretical value of
445 Torgersen et al. (1995). Torgersen et al. (1995) considered this concentration to be a good proxy
446 for a hot spot source. Indeed plume-derived melts should represent the non-degassed lower mantle,
447 which is expected to have preserved more volatiles than the upper degassed mantle, in the canonical
448 view of an Earth' mantle layered reservoir based on noble gas systematics (e.g., Allègre et al.,
449 1995). However, OIBs are constantly depleted in helium compared to the MORBs (e.g., Moreira
450 and Kurz, 2013). This contradiction is known as the “mantle helium paradox” (e.g., Hopp and
451 Trieloff, 2008). There are several models to explain this paradox (e.g., Hopp and Trieloff, 2008;
452 Gonnermann and Mukhopadhyay, 2007) but basically it is assumed that plume melts undergone a
453 non-equilibrium degassing elemental fractionation during their ascent to the surface (Gonnermann
454 and Mukhopadhyay, 2007). This model is based on the effect of CO_2 and H_2O contents in the melts
455 which can drastically reduce the solubility of noble gases. OIB melts would be enriched in CO_2
456 compared to MORB ones, decreasing substantially the solubility of He which will tend to be largely
457 degassed. Gonnermann and Mukhopadhyday (2007) constrained the He content in the pre-
458 degassed melt to be $3 \times 10^{-4} \text{ cm}^3/\text{g}_{\text{rock}}$, suggesting that more than 99 % of helium is lost during non-
459 equilibrium magmatic degassing. This ${}^4\text{He}$ content is 2 orders of magnitude higher than that
460 expected for the MORB magma source ($3 \pm 2 \times 10^{-6} \text{ cm}^3\text{STP/g}_{\text{rock}}$; Moreira and Kurz, 2013).

461 All parameters of Eqn. (6) with assumed or measured uncertainties are summarized in Table
462 3. An alternative way to calculate ${}^4\text{He}_{\text{initial}}$ is from rock analyses (Table 2). However, this is limited
463 by the diffusion and loss of He from the rocks and the distribution of U and Th in the rock matrix
464 and its effect on the dilution of the mantle signal, which is likely preserved in U and Th-poor
465 minerals, such as in clinopyroxene.

466 The initial ${}^3\text{He}/{}^4\text{He}$ ratio expected in the fossil gas-rich magma source in the region
467 ($R_{\text{mtl}}/R_{\text{a}_{\text{initial}}}$; Eqn. (6)) has been evaluated, together with its total uncertainty, using a Monte Carlo
468 simulation. Monte Carlo simulations make use of the uncertainty domains for all given variables
469 (Table 3) in a specific model, here the magma aging model explicated in Eqn. (6). By running the
470 algorithm multiple times –a hundred thousand times here – with random sampling of those
471 variables, individual output data can be generated. The probability distribution of the outputs is
472 usually reported as a Gaussian curve, and the average and standard deviation thereof can be
473 extrapolated.

474 Monte Carlo simulations were carried out using the NIST Uncertainty Machine
475 (<https://uncertainty.nist.gov/>), which is a web-based software application to evaluate the
476 measurement uncertainty associated with an output quantity defined by a measurement model of
477 the form $y = f(x_1, \dots, x_n)$ (Lafarge and Possolo, 2015). Input parameters (Table 3) are introduced
478 online, the number of iterations is selected, and the measurement model (Eqn. 6) is compiled in an
479 “R” script.

480 Two simulations were carried out, representing two extreme cases (Table 3). The first one
481 considers the progressive aging of a magma source equal to the alnöite, which emplaced between
482 107 and 114 Ma ago. The second one considers the progressive aging of an average magma source

483 of gabbro-sienitic composition, equal to the intrusions of the Montereian Hills and emplaced
484 124 ± 1 Ma ago. A ${}^4\text{He}_{\text{initial}}$ content of $6.7 \pm (10\%) \times 10^{-6} \text{ cm}^3\text{STP/g}_{\text{rock}}$ is assumed, equivalent to the
485 least degassed OIBs. We could consider an even more degassed OIBs source, however, for ${}^4\text{He}_{\text{initial}}$
486 contents $\leq 10^{-7} \text{ cm}^3\text{STP/g}_{\text{rock}}$, the initial helium isotopic ratios would be rapidly diluted by *in situ*
487 radiogenic ${}^4\text{He}$ production on a timescale of 10 Ma or even less, completely masking any mantle
488 helium record in present-day volcanics.

489 The outputs of the Monte Carlo simulations have been represented in two different ways.
490 The first, reported in Fig. 5 is as the resulting Gaussian curves for alnöite and for the Montereian
491 Hills intrusions (MH) simulations. The first simulation for the alnöite resulted in a $R_{\text{mtl}}/R_{\text{a}_{\text{initial}}}$ of
492 21 ± 10 (2σ), while that for the gabbro-sienitic MH resulted in a $R_{\text{mtl}}/R_{\text{a}_{\text{initial}}}$ of 33 ± 28 (2σ). The
493 larger uncertainty for the second simulation is due to the larger error assumed on the $J_{4\text{He}}$ (Eqn. 5
494 and Table 2) parameter. The output probability distributions for the two simulations have been
495 compared with the frequency histogram of the measured ${}^3\text{He}/{}^4\text{He}$ ratios in 401 samples of OIB
496 glasses (Fig. 5). The latter data was extracted from the global dataset of noble gas abundances and
497 isotopic ratios in Volcanic Systems – USGS-NoGaDat, compiled by Abedini et al. (2006). Present-
498 day plume-sourced helium in OIB glass phase has an average R/R_{a} of 15.8 ± 8.6 (1σ) (Fig. 5).

499 Results of the Monte Carlo simulations have also been reported in a clearer way (Fig. 6),
500 where the progressive “aging” of the initial helium isotopic signature of the emplaced magma to
501 the present-day measured value is reported as a curve, together with its uncertainty envelope
502 calculated from the Monte Carlo simulations. For the two simulations, up to 96% of the results
503 show $(R_{\text{mtl}}/R_{\text{a}})_{\text{initial}}$ higher than that expected for a lithospheric magma source, either the convective
504 upper mantle ($R=8 \pm 1 R_{\text{a}}$; Allègre et al., 1995) or the SCLM ($R/R_{\text{a}}=6.1 \pm 0.9$; Gautheron and

505 Moreira, 2002) (Fig. 6a). In this simulation, the mantle $^3\text{He}/^4\text{He}$ ratio is assumed to have been
506 constant in the last 120Ma, which is what is expected from numerical simulations of the helium
507 isotopic evolution in mantle reservoirs (e.g., Porcelli and Elliot, 2008).

508 These two simulations suggest that a mantle plume was possibly the source of the fossil
509 helium presently found in the groundwater of the southern Quebec, which may be related to the
510 New England hotspot.

511 If the previous Monte Carlo simulations were carried out using an undegassed OIB source
512 $^4\text{He}_{\text{initial}}$ content of $3.14 \pm (10\%) \times 10^{-4} \text{ cm}^3\text{STP/g}_{\text{rock}}$ (Gonnermann and Mukhopadhyay, 2007), then
513 the resulting $(R_{\text{mtl}}/R_{\text{a}})_{\text{initial}}$ will be very low from $1.8 \pm 0.2R_{\text{a}}$ (alnoïte) to $2.4 \pm 0.4R_{\text{a}}$ (Monteregian
514 Hills). These values are even lower than those expected for a MORB or the SCLM-type magma
515 source. Further, it would be very difficult to explain how the plume-generated melts beneath the
516 Monteregian Hills could have preserved the pristine helium content when globally, OIBs and
517 continental basalts show constanly different degrees of degassing (e.g., Moreira and Kurz, 2013;
518 Kurz et al., 2009). It is worth noting that these two last simulations cannot represent a local MORB-
519 type or SCLM source, because expected ^4He content in MORBs are 2 orders of magnitude lower
520 than those used for these simulations (Moreira and Kurz, 2013).

521

522 **6. Conclusions**

523 This study highlights the possibility of long-term lithospheric storage of primordial noble
524 gases from past magmatic events. The occurrence of mantle-derived helium in the eastern North
525 America craton, first suggested by Torgersen et al. (1995) and then argued by this study, should be
526 further investigated. Analyses of groundwater around intrusions related to the New England

527 hotspot, such as the White Mountains, should be carried out to confirm the preservation of this
528 mantle memory. Such analyses would also better constrain the hydrogeology of fractured igneous
529 intrusions, which is still poorly documented and understood. The processes leading to the storage
530 of this mantle memory and to its release into groundwater remain to be identified. Recently, Méjean
531 et al. (2017) showed that the release of helium into groundwater could be episodic and related to
532 tectonic events that increase the local fracturing of the rocks. During these events, the release of
533 mantle helium preserved in high-temperature magmatic mineral phases, such as olivine or
534 clinopyroxene, could be substantial. It is then diluted over time by the progressive recharge of the
535 aquifer by freshwater (containing atmospheric helium) and the local accumulation of radiogenic
536 helium. This study suggests that groundwater could preserve a record of mantle helium on a
537 timescale of hundreds of millions of years, and may even provide a more faithful record of the
538 primary signature compared to the magmatic rocks from which the He is ultimately derived.

539

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693

694 **Figure Captions**

695 **Figure 1.** Map of the St. Lawrence Lowlands, including the study area of the Vaudreuil-Soulanges
696 watershed (VS) and the intrusions of the Cretaceous Monteregian Hills (MH). The Bécancour
697 watershed (BEC) and the Nicolet and lower Saint-Francois River watershed (NSF) are also
698 indicated. The percentage of mantle helium ($f_{\text{mtl}}\%$) measured in each watershed is provided.

699 **Figure 2.** a) Map of the sampled groundwater and rocks in the VS watershed, with outcrops of the
700 Cretaceous Oka carbonatite complex and the Île Cadieux alnoïte indicated, together with the
701 Grenvillian Mount Rigaud. b) Map of the sampled groundwater and rocks of the Monteregian Hills.

702 **Figure 3.** Distribution of R/Ra values measured in bulk rocks and mineral separates from
703 Monteregian Hills intrusions by Sasada et al. (1997) and in this study. The range of mantle-derived
704 $^3\text{He}/^4\text{He}$ measured in groundwater of the St. Lawrence Lowlands basin is also reported for
705 comparison.

706 **Figure 4.** (a) Theoretical Weise-type diagram showing the main helium sources, namely the
707 atmosphere, the crust and the mantle, and how the helium isotopic and elemental compositions
708 vary by adding tritiogenic He (light blue arrow), crustal He (yellow arrow), and mantle He (orange
709 arrow) to an atmospheric helium source. A mixing line between a hypothetical tritiogenic helium
710 end-member and sample WSB1 is reported. The intercept of this mixing line with the Y-axis gives
711 the fraction of mantle He in the mixture. (b) Weise diagram of the measured helium ratios. The
712 black line represents mixing between water at recharge (R_{eq}) with water having accumulated crustal
713 radiogenic helium (R_{rad}). Dashed lines through samples represent mixing between modern water of

714 different initial tritium contents mixed with water containing terrigenous helium (mantle and crustal
715 helium).

716 **Figure 5.** Probability distributions of the initial $^3\text{He}/^4\text{He}$ ratio (R_{mtl}/R_a ; Eqn. 6) of the emplaced
717 magma source (alnoïte in the Vadreuil-Soulanges (VS) area and gabbro-sienitic for the
718 Monteregian Hills (MH)), obtained from the Monte Carlo method. These distributions are plotted
719 against the frequency distribution of the R/R_a measured in 401 OIB glasses (data from Abedini et
720 al., 2006). SD = standard deviation (1σ); MAD= mean absolute deviation.

721 **Figure 6.** Evolution curves (and uncertainty envelope) of the (R_{mtl}/R_a) as a function of time elapsed
722 since the emplacement of the alnoïte intrusions (a) in the VS region 110 ± 3.5 Ma; and (b) in the
723 sampled Monteregian Hills (124 Ma), as calculated by Monte Carlo simulations.

724

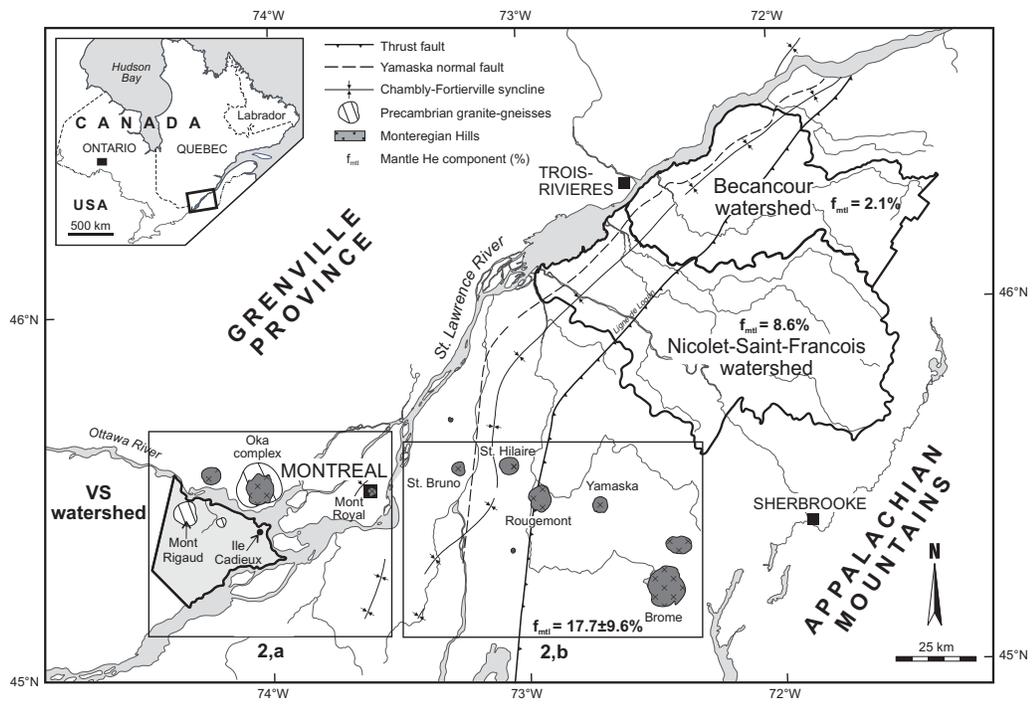


Figure 1
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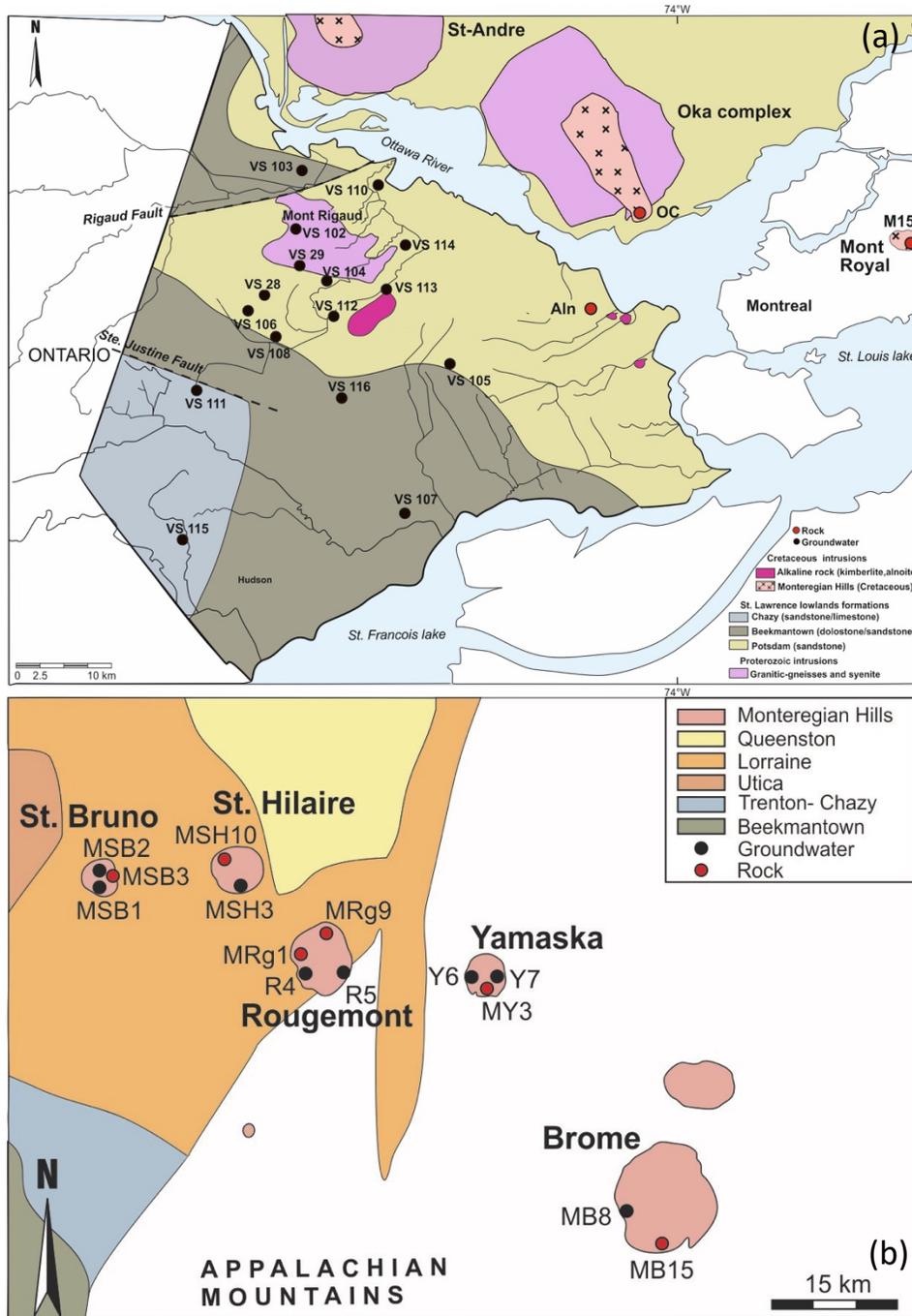


Figure 2 a,b
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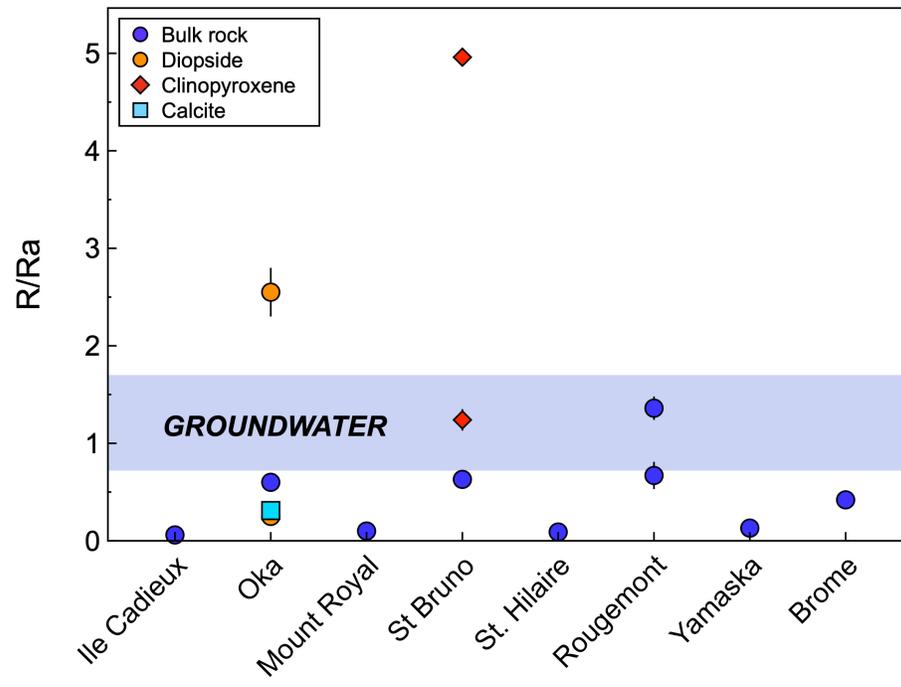


Figure 3
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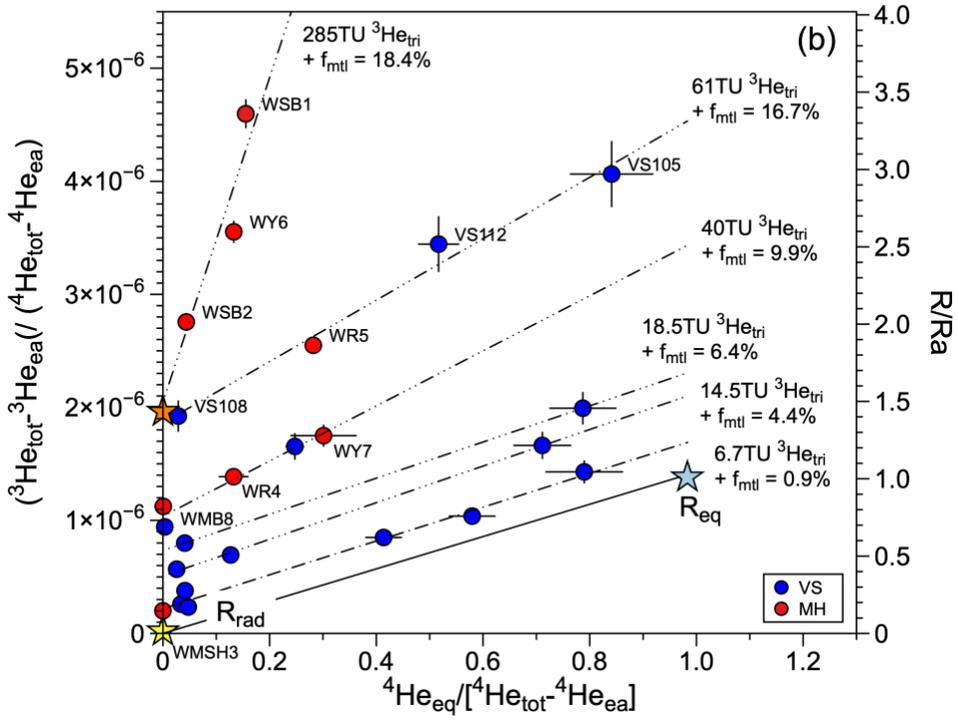
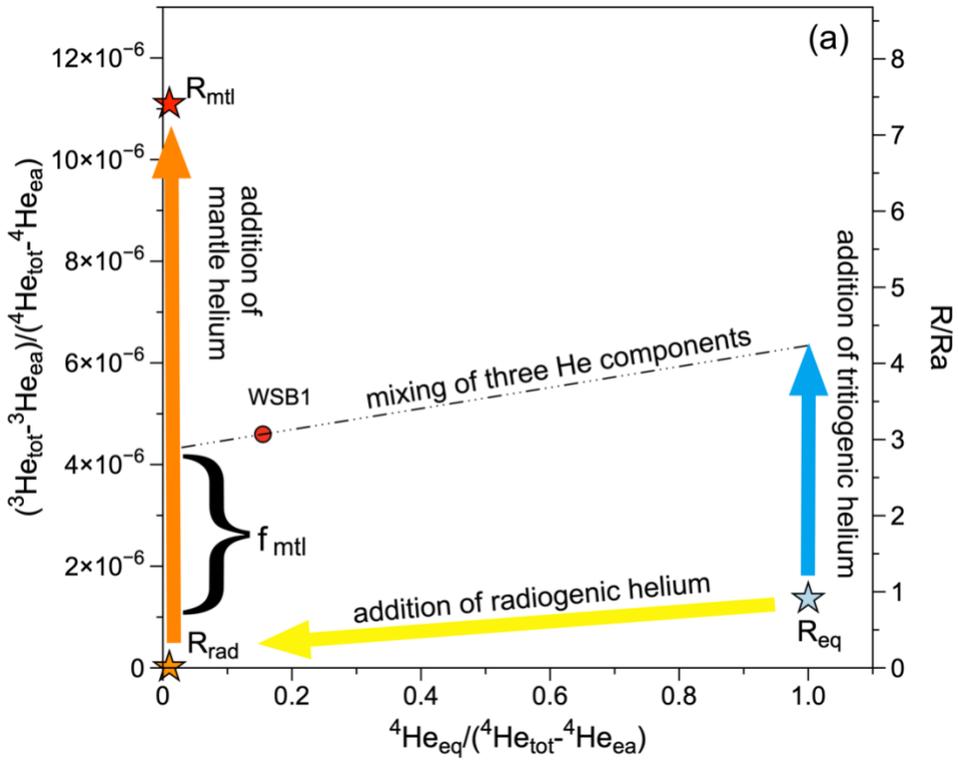


Figure 4 a,b

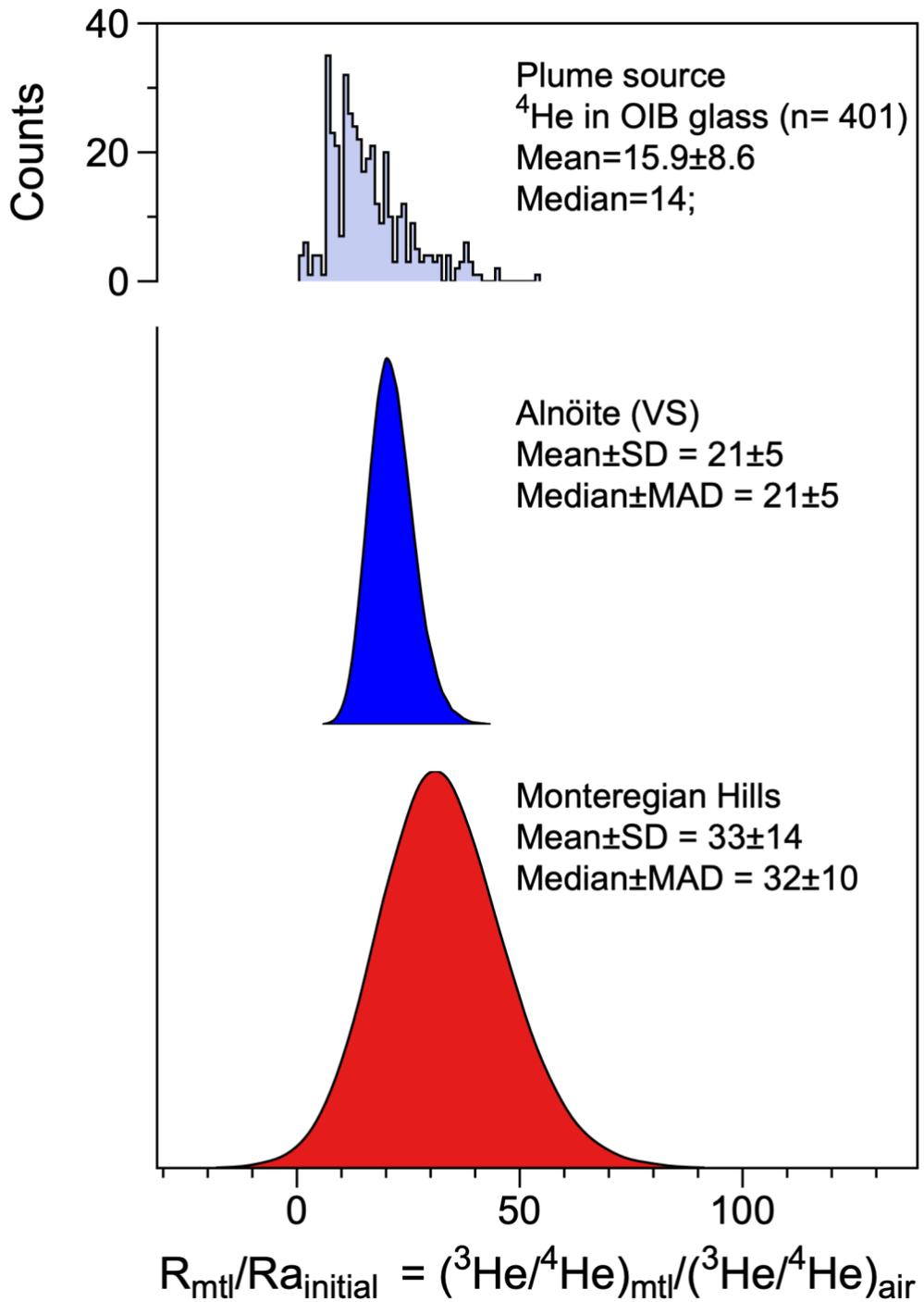


Figure 5
 EPSL-D-19-01096R1

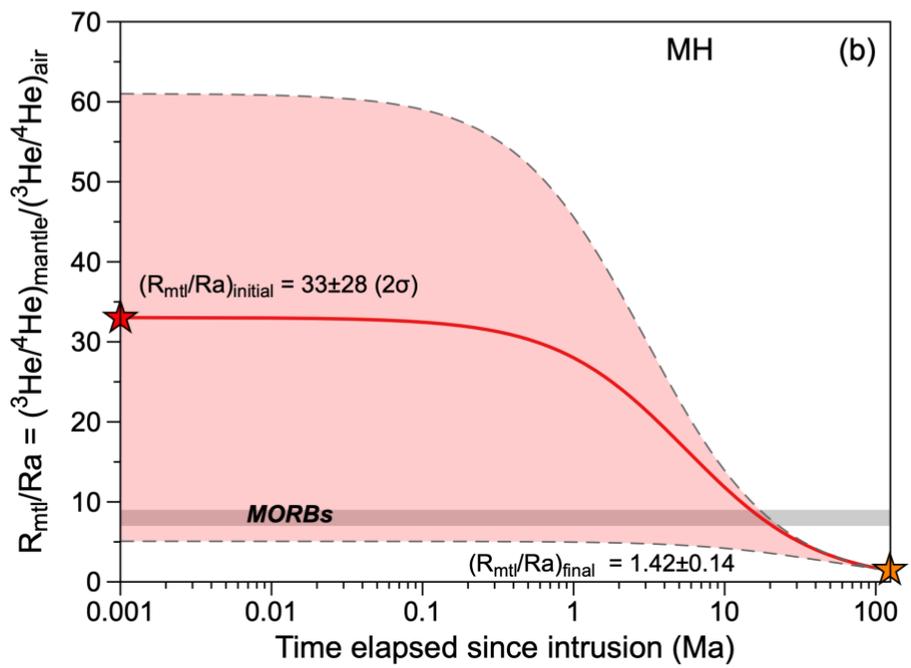
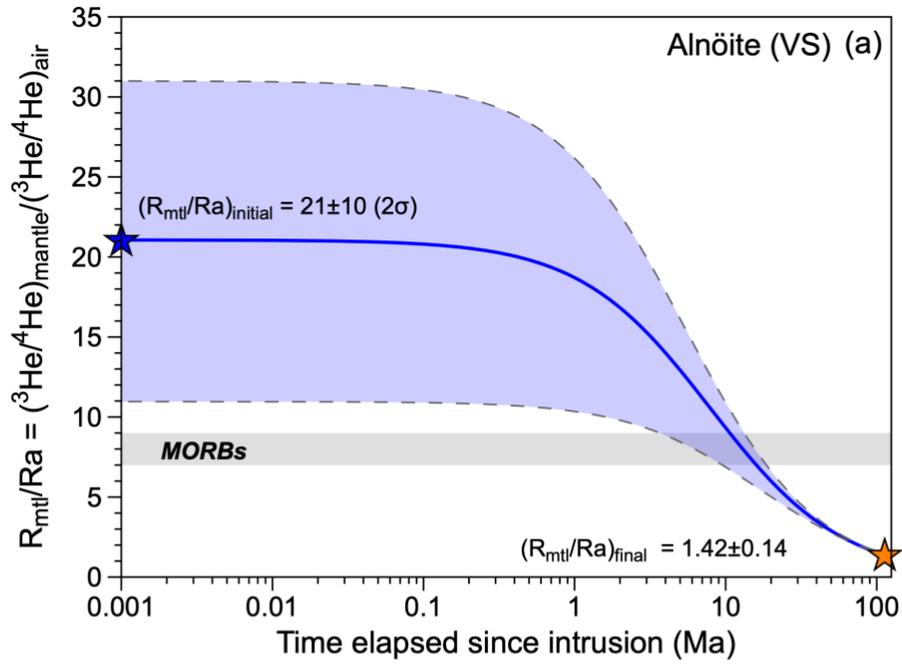


Figure 6 a,b

EPSL-D-19-01096R1

Table 1. Helium and neon elemental and isotopic data of sampled groundwater.

Sample	Type of well	Location	Latitude	Longitude	⁴ He		²⁰ Ne		R/Ra		⁴ He/ ²⁰ Ne		ΔNe% (ea)
					cm ³ STP/g x 10 ⁻⁸	±	cm ³ STP/g x 10 ⁻⁷	±		±		±	
					UTM NAD83								
VS102	Municipal	Vaudreuil-Soulanges	-74.31279	45.44997	9.28	4.64	2.87	0.14	1.02	0.01	0.32	0.16	54.48
VS103	Municipal	Vaudreuil-Soulanges	-74.31123	45.48339	116.10	5.81	2.11	0.10	0.58	0.01	5.51	0.39	12.83
VS104*	Municipal	Vaudreuil-Soulanges	-74.28452	45.42295	6.64	0.33	1.72	0.09	1.20	0.02	0.39	0.03	-8.02
VS105	Observation	Vaudreuil-Soulanges	-74.181861	45.381763	98.81	0.44	2.83	0.14	2.23	0.03	0.31	0.02	51.34
VS106	Observation	Vaudreuil-Soulanges	-74.349212	45.401767	19.46	0.97	1.95	0.10	1.19	0.01	1.00	0.07	4.28
VS107	Observation	Vaudreuil-Soulanges	-74.208624	45.296467	7.46	0.37	2.28	0.11	1.35	0.02	0.33	0.02	21.93
VS108	Private	Vaudreuil-Soulanges	-74.32428	45.3881	166.64	8.33	2.79	0.14	1.38	0.02	5.97	0.42	49.20
VS110	Private	Vaudreuil-Soulanges	-74.25092	45.47927	9.26	0.46	2.17	0.11	0.78	0.01	0.43	0.03	16.04
VS111*	Private	Vaudreuil-Soulanges	-74.38018	45.35587	139.09	6.95	1.75	0.09	0.19	0.01	7.96	0.56	-6.42
VS112	Private	Vaudreuil-Soulanges	-74.2761	45.40286	10.019	0.51	2.16	0.11	2.33	0.03	0.47	0.03	15.51
VS113*	Private	Vaudreuil-Soulanges	-74.23643	45.42119	183.57	9.18	1.53	0.09	0.41	0.01	11.96	0.94	-18.18
VS114*	Private	Vaudreuil-Soulanges	-74.22512	45.44675	1323	68.01	1.79	0.09	0.68	0.01	74.25	5.32	-4.28
VS115	Private	Vaudreuil-Soulanges	-74.38136	45.27199	100	5.03	2.24	0.11	0.18	0.004	4.48	0.32	19.79
VS116	Private	Vaudreuil-Soulanges	-74.26511	45.35669	39.39	1.97	2.54	0.13	0.53	0.01	1.55	0.11	35.83
VS28	Private	Vaudreuil-Soulanges	-74.33418	45.41088	15.19	0.76	3.01	0.15	0.71	0.01	0.50	0.04	60.96
VS29	Private	Vaudreuil-Soulanges	-74.30626	45.42946	139.59	9.41	2.1	0.07	0.28	0.01	6.65	0.50	12.30
WSB1*	Private	MH, St-Bruno	-73.3241974	45.5431257	30.43	0.29	1.76	0.02	3.32	0.08	1.73	0.02	-5.88
WSB2*	Private	MH, St-Bruno	-73.3249577	45.5441935	108.20	1.01	0.67	0.01	1.99	0.05	16.18	0.21	-64.17
WMSH3	Private	MH, St-Hilaire	-73.156288	45.537288	7857827	73675	15780	147	0.15	0.01	49.80	0.66	-
WMSH3 (corrected)					93.12	0.87	1.87	0.02					-
WR4*	Private	MH, Rougemont	-73.080321	45.4633378	35.73	0.33	0.52	0.01	1.00	0.03	6.81	0.09	-72.19
WR5	Private	MH, Rougemont	-73.0336287	45.4605797	17.39	0.16	2.02	0.02	1.81	0.05	0.86	0.01	8.02
WY6	Private	MH, Yamaska	-72.88344	45.46796	38.16	0.36	2.63	0.02	2.46	0.06	1.46	0.02	40.64
WY7*	Private	MH, Yamaska	-72.86829	45.46526	15.72	0.15	0.54	0.01	1.26	0.07	2.92	0.04	-71.12
WMB8	Private	MH, Brome	-72.694374	45.262202	12685	119	1435	13	0.88	0.03	0.88	0.01	-
WMB8 (corrected)					1.65	0.94	1.88	0.02					-

VS = Vaudreuil Soulanges watershed; MH = Monteregian Hills

*⁴He was corrected for degassing when ²⁰Ne is lower than the atmospheric value (solubility of Ne at 9.8°C is 1.87 x 10⁻⁷ cm³STP/g)ΔNe (%), given as ([Ne]_{measured}/[Ne]_{atmospheric equilibrium} - 1) x 100, and represents the total amount of excess air

Table 2. Measured helium isotopic composition, calculated helium ratios and He production rates in

Sample	Location	Rock type	Phase	R/Ra	±	[⁴ He]	±
						cm ³ STP/g x 10 ⁻⁵	
R-Aln	Ile Cadieux	Alnöite	Bulk	0.06	0.02	0.70	0.07
R-OC	Oka Complex	Carbonatite	Bulk	0.60	0.01	1.28	0.12
OKDI*	Oka Complex	Carbonatite	Diopside	0.25	0.07	0.82	0.08
OKDI*	Oka Complex	Carbonatite	Diopside	2.55	0.25	0.13	0.01
OKCA*	Oka Complex	Carbonatite	Calcite	0.31	0.09	1.16	0.12
R-MR15	Mont Royal	Essexite	Bulk	0.10	0.03	0.95	0.09
R-MSB3	St. Bruno	Felsic gabbro	Bulk	0.62	0.06	0.55	0.05
R-MSB3	St. Bruno	Felsic gabbro	Clx-pyroxene	1.23	0.10	0.48	0.04
R-MSB3	St. Bruno	Felsic gabbro	Clx-pyroxene	4.96	0.01	0.07	0.01
R-MB15	Mont Brome	Gabbro	Bulk	0.43	0.03	0.41	0.04
R-MRg1	Mont Rougemont	Banded gabbro	Bulk	0.67	0.13	0.07	0.01
R-MRg9	Mont Rougemont	Gabbro	Bulk	1.36	0.11	0.41	0.04
R-MSH10	Mont St. Hilaire	Gabbro	Bulk	0.09	0.01	2.28	0.22
R-MY3	Mont Yamaska	Gabbro	Bulk	0.13	0.01	1.92	0.19

* Data from Sasada et al. (1997)

**Rin/Ra is the resulting helium isotopic ratio from in situ production of ³He and ⁴He.

1 Monterey Hills' rocks and minerals.

[U]	[Th]	[Li]	$J^3\text{He}$	\pm	$J^4\text{He}$	\pm	R_{in}/Ra^{**}
ppm	ppm	ppm	$\text{cm}^3\text{STP/g}_{rock}/\text{y}$		$\text{cm}^3\text{STP/g}_{rock}/\text{y}$		
			$r \times 10^{-20}$		$r \times 10^{-13}$		
3.3	15.4	15	1.78	0.36	8.41	1.68	0.015
9.9	40.9	15	0.63	0.13	23.56	4.71	0.002
n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.
n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.
n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.
4.5	9.6	37	1.58	0.32	8.48	1.7	0.013
3.9	9.4	28	1.04	0.21	7.35	1.5	0.010
n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.
n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.	n.d.
12.4	20.1	28	2.89	0.58	20.54	4.1	0.010
4.5	8.8	28	1.11	0.22	7.89	1.58	0.010
4.1	9.4	28	1.07	0.21	7.59	1.5	0.010
10.0	19.6	28	2.47	0.50	17.64	3.5	0.010
6.6	17.7	28	1.82	0.36	13.0	2.6	0.010

Table 3. Summary of input and output parameters for Monte Carlo simulations.

Input parameters	units	Alnoite	\pm	MH	\pm
J ₄ He	$\times 10^{-13}$ ccSTP/g _{rock} /yr	8.41	1.68	11.78	5.01
(R _{mtl} /Ra) _{final}		1.42	0.14	1.42	0.14
⁴ He initial	$\times 10^{-6}$ ccSTP/g	6.7	0.7	6.7	0.7
Age since intrusion	Ma	110.5	3.5	124	1
Output parameter :					
(R _{mtl} /Ra) _{initial}	Average*	21	5	33	14
(R _{mtl} /Ra) _{initial}	Median**	21	5	32	10

Monte Carlo simulation results after 1×10^5 iterations.

* Uncertainty on the average value is standard deviation (1σ)

** Uncertainty on the median is the mean absolute deviation (MAD)